

## Orbital Periods of a Satellite

In orbital mechanics we are concerned with the following orbital periods of a satellite:

- *Keplerian* - the unperturbed or two-body period
- *Nodal* - the time interval from one ascending (or descending) node to the next
- *Anomalistic* - the time interval from one perigee to the next
- *Sidereal* - the time interval from one value of argument of latitude to the next identical value

This *Numerit* program (`period1`) can be used to estimate these orbital periods. Each algorithm is based on the *osculating* orbital elements of a satellite and includes the perturbation due to  $J_2$ .

The Keplerian or unperturbed orbital period of a satellite is given by

$$t_k = 2\mathbf{p}\sqrt{\frac{a^3}{\mathbf{m}}} \quad (1)$$

The anomalistic period of a satellite based on osculating orbital elements at perigee is given by

$$t_a = 2\mathbf{p}\sqrt{\frac{a^3}{\mathbf{m}}} \left\{ 1 - \frac{3J_2 r_{eq}^2}{2a^2(1-e^2)^3} (1 - 3\sin^2 i \sin^2 \mathbf{w}) \right\} \quad (2)$$

The nodal period of a satellite based on osculating orbital elements of the satellite at the ascending node is

$$t_n = 2\mathbf{p}\sqrt{\frac{a^3}{\mathbf{m}}} \left\{ 1 - \frac{3J_2(4 - 5\sin^2 i)}{4\left(\frac{a}{r_{eq}}\right)^2 \sqrt{1-e^2}(1+e\cos\mathbf{w})^2} - \frac{3J_2(1+e\cos\mathbf{w})^3}{2\left(\frac{a}{r_{eq}}\right)^2 (1-e^2)^3} \right\} \quad (3)$$

The sidereal period of a satellite based on osculating orbital elements at perigee of the orbit is given by

$$t_s = 2\mathbf{p}\sqrt{\frac{a^3}{\mathbf{m}}} \left\{ 1 - \frac{3}{2}J_2\left(\frac{ar_{eq}^2}{r^3}\right)(1 - 3\sin^2 i \sin^2 u) - \frac{3}{4}J_2\left(\frac{r_{eq}}{a}\right)^2 \frac{1}{\sqrt{1-e^2}} \left[ \frac{4 - 5\sin^2 i}{(1+e\cos\mathbf{u})} \right] \right\} \quad (4)$$

## Orbital Mechanics with Numerit

In these equations

$a$  = semimajor axis  
 $e$  = orbital eccentricity  
 $i$  = orbital inclination  
 $w$  = argument of perigee  
 $u$  = true anomaly  
 $u$  = argument of latitude =  $w + u$   
 $r$  = geocentric distance  
 $r_{eq}$  = equatorial radius of the Earth  
 $m$  = gravitational constant of the Earth  
 $J_2$  = oblateness gravity coefficient

The following is a typical draft output created with this program.

```
program period1
< orbital periods - j2 analytic solution >

semimajor axis      8000 kilometers
eccentricity        0.015
inclination         28.5 degrees
argument of perigee 270 degrees
true anomaly        30 degrees

Keplerian period    118.68468 minutes
nodal period        118.38678 minutes
anomalous period    118.64405 minutes
sidereal period     118.45169 minutes
```